

# Approaches for Estimating Radioxenon Background Variations, Anomalies, and Explosion Signals in Modeled and Measurement Data

## *Presenter*

Donald Lucas\*



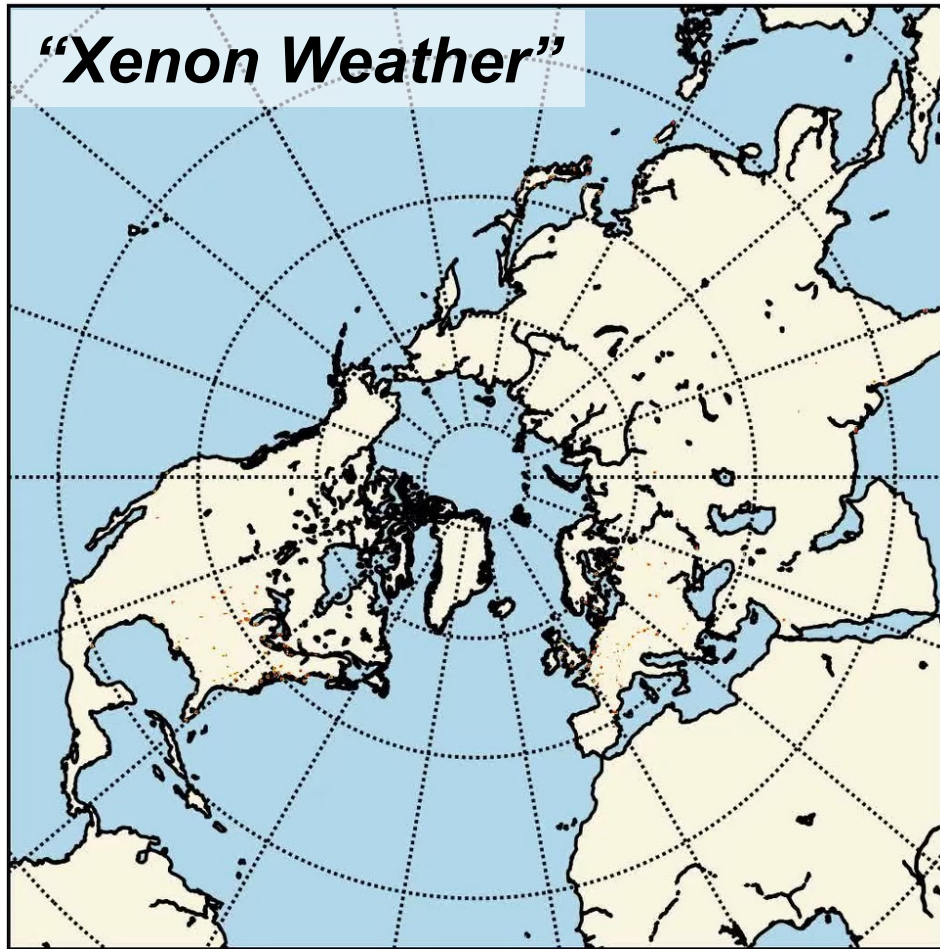
## *Team* (alphabetical order)

Ted Bowyer<sup>+</sup>, Paul Eslinger<sup>+</sup>, Nipun Gunawardena<sup>\*</sup>, Lee Glascoe<sup>\*</sup>, Donald Lucas<sup>\*</sup>, John Lucas<sup>^</sup>, Lucas Reilly<sup>^</sup>, John Roberts<sup>^</sup>, and Ramesh Sarathi<sup>+</sup>



# Background Radioxenon is Highly Variable in Space and Time

2014Jul01-00



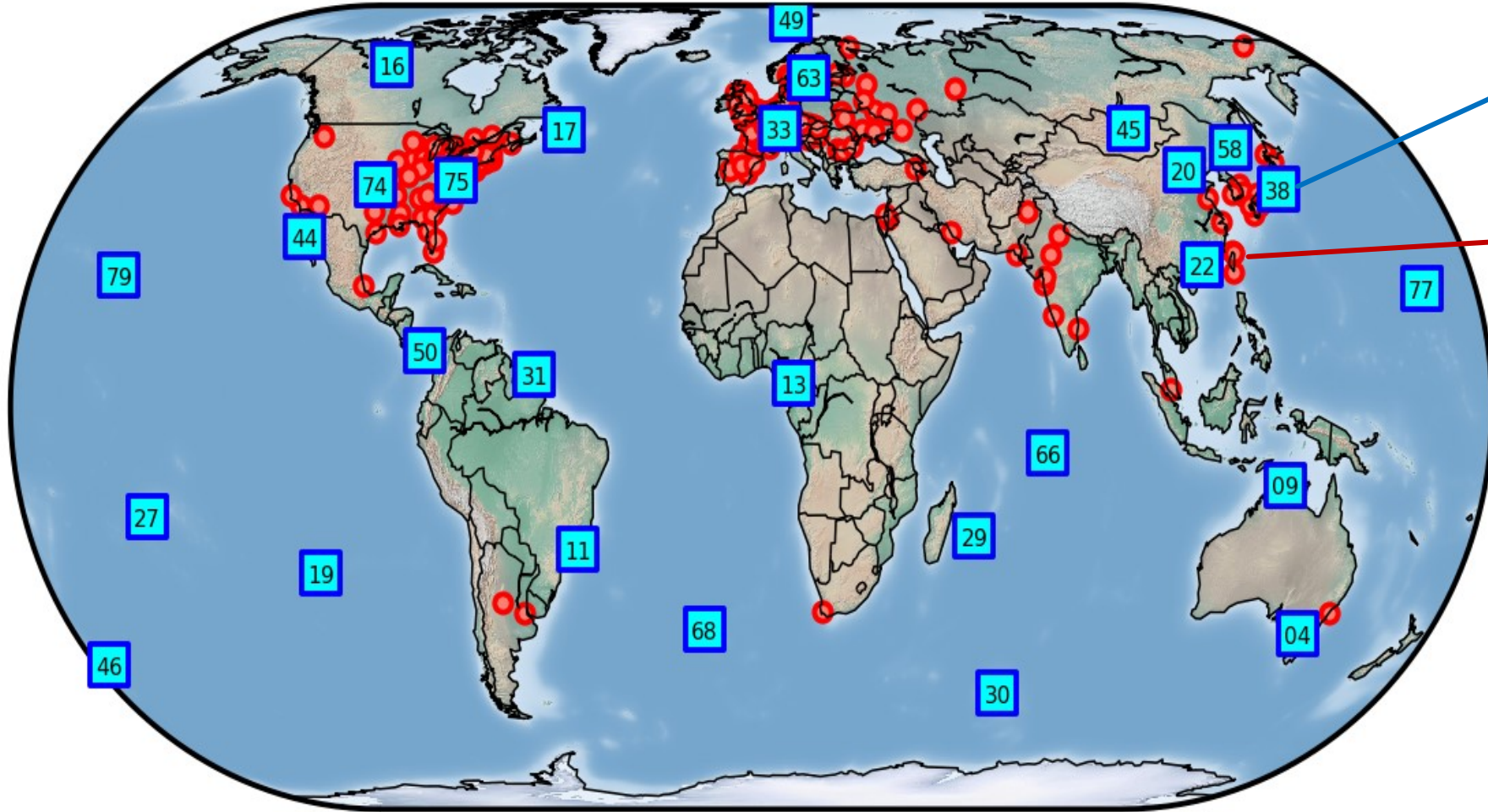
Movie of Xe-133 released from 200 facilities on 2014 July 01 and tracked for two weeks. Colors show near-surface logarithmic activity concentrations.

Extracting nuclear test signals from the radioxenon background is like finding a needle in a haystack.



Advances in modeling and algorithms may help find the needle.

# Background Radioxenon is Highly Variable in Space and Time

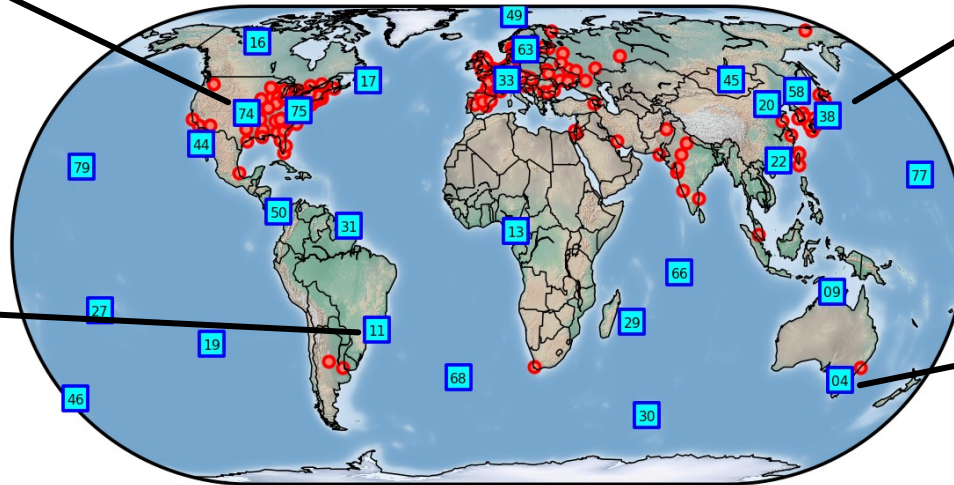
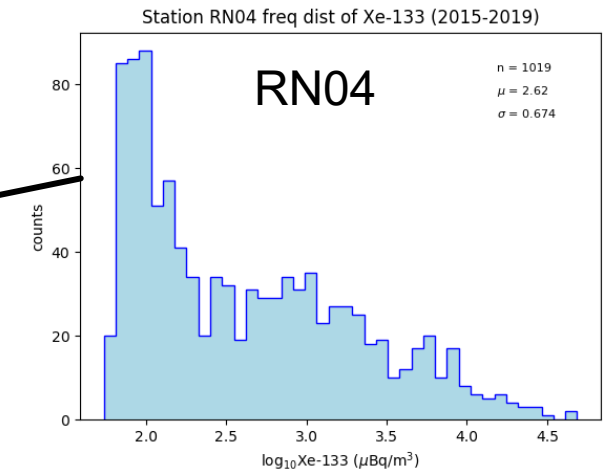
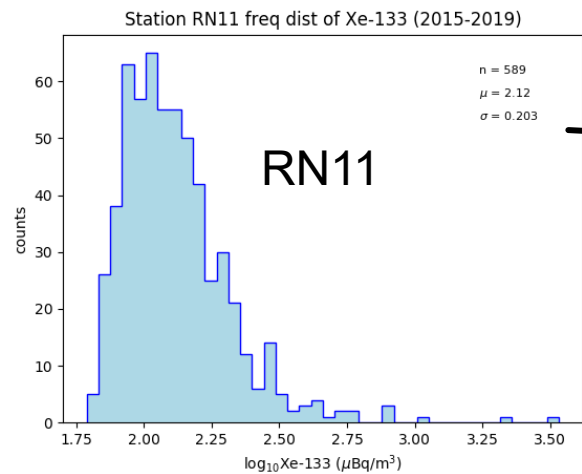
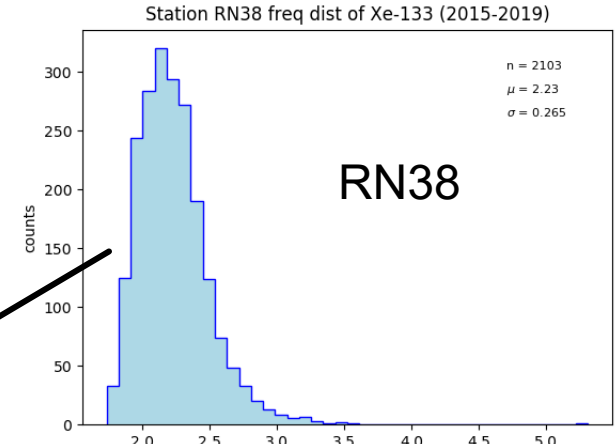
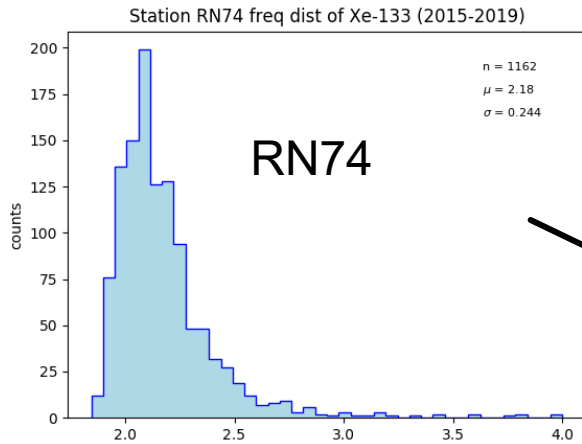


RN stations

Potential background xenon sources (nuclear power plants, research reactors, and medical isotope facilities)

# Background Radioxenon is Highly Variable in Space and Time

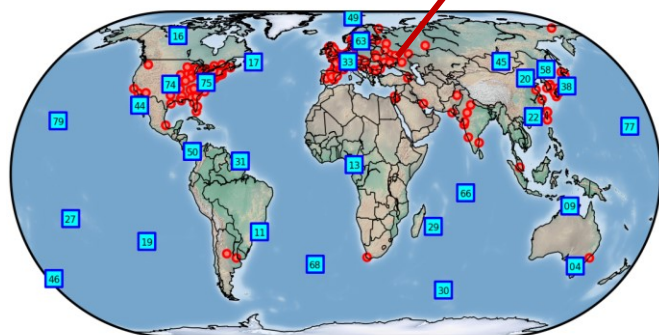
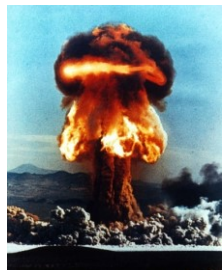
**Radioxenon variability** at IMS stations depends on proximity to sources, changes in meteorology, topography, and land properties, and other factors.



Frequency distributions of Xe-133 over the period 2015-2019.

# Atmospheric Models Can Be Used to Estimate Background Xenon

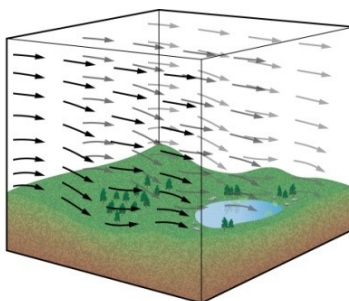
## Xenon Sources



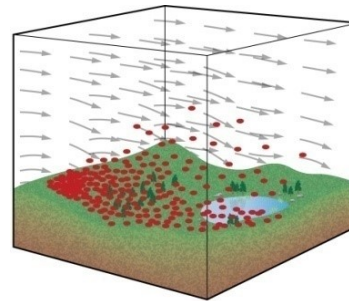
Xenon emission rates at facilities are typically not measured and can represent a large source of model uncertainty.

## Atmospheric Models

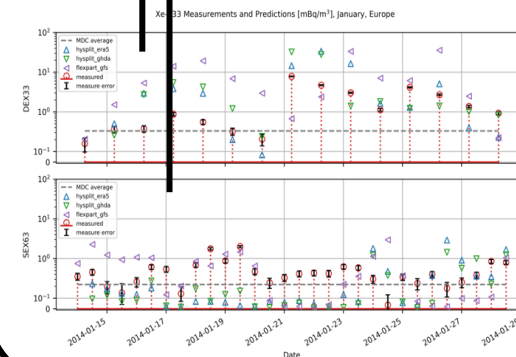
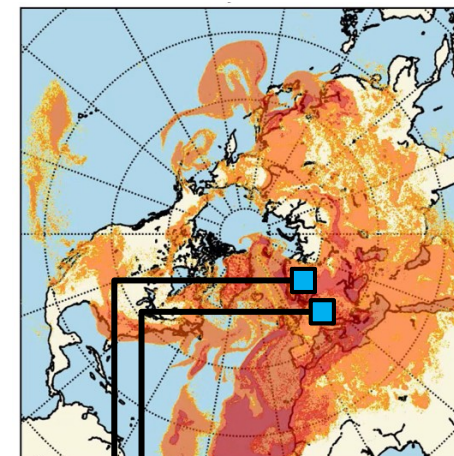
### Meteorology



### Atmospheric transport



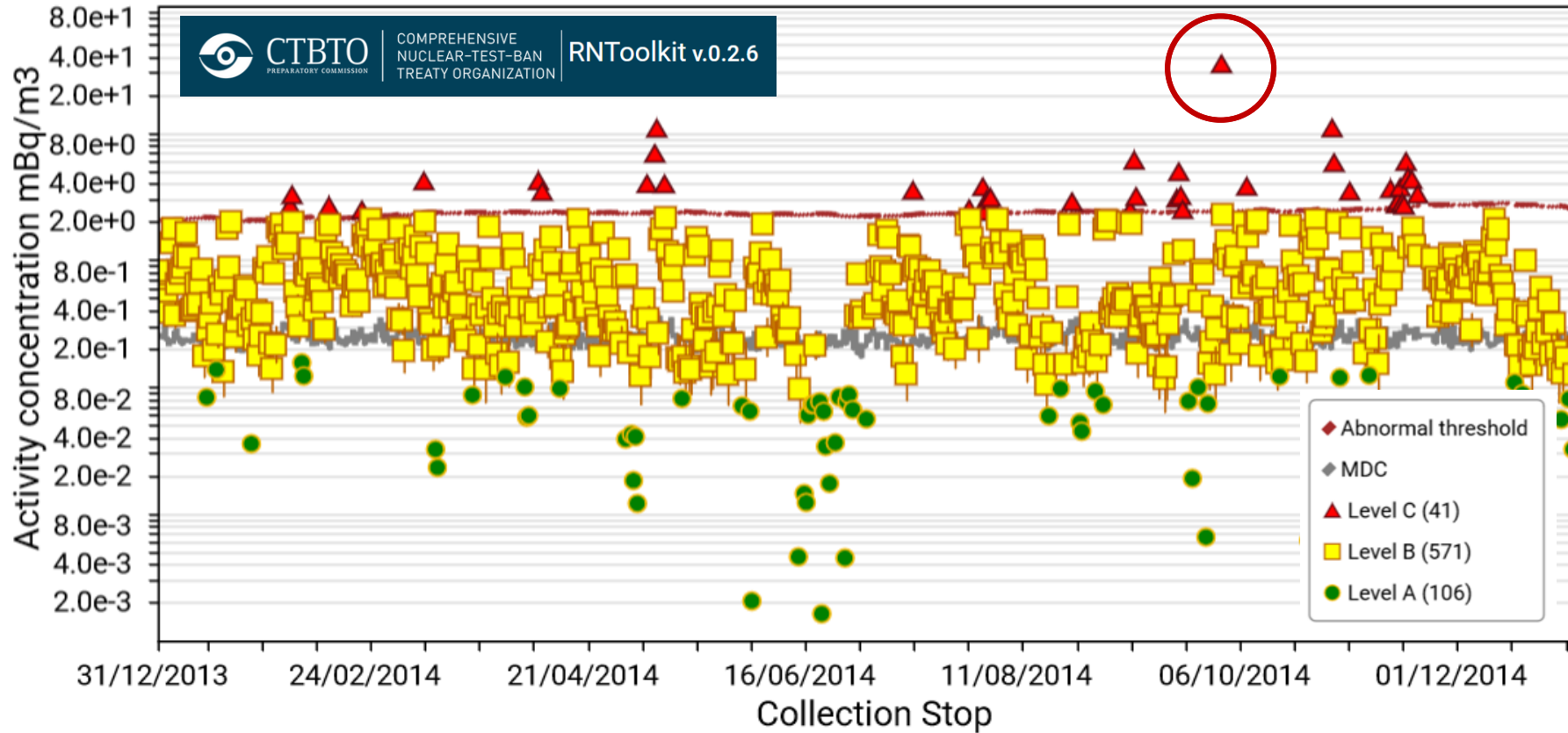
## Xenon Signals



Inverse Modeling

# Identifying Xenon Anomalies in Measurements

Xe-133 history at SEX63 - Long term - Interactive analysis



To distinguish explosion signals from background sources, it is important to quantify the size and frequency of xenon anomalies.

# Identifying Xenon Anomalies in Measurements

Anomalies can occur in multiple dimensions

One or more Xe isotopes

One or more IMS stations

Quantile Scores

Empirical and easy to compute in one dimension, but challenging in higher dimensions

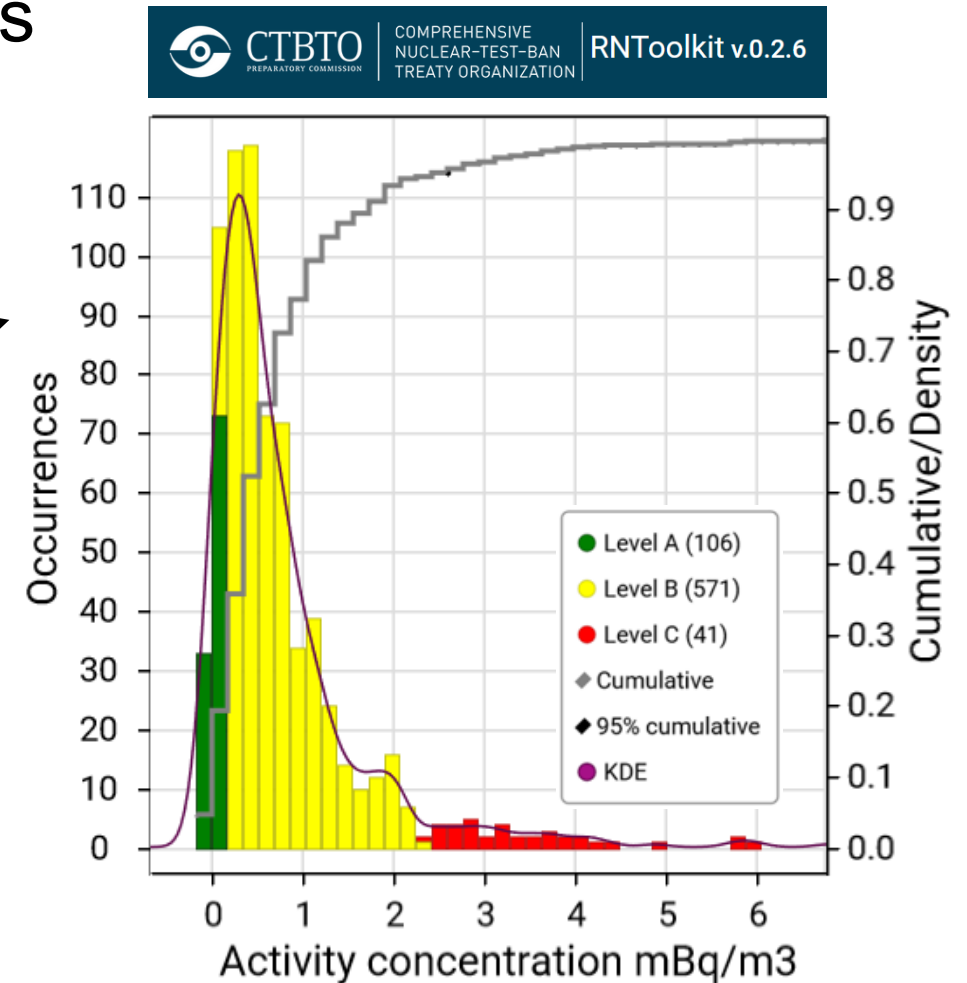
Outlier/Novelty Detection Algorithms

Time series methods

Machine learning approaches

Local Outlier Factor

Random Isolation Forest



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- Time series methods

- Machine learning approaches

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  - Random Isolation Forest

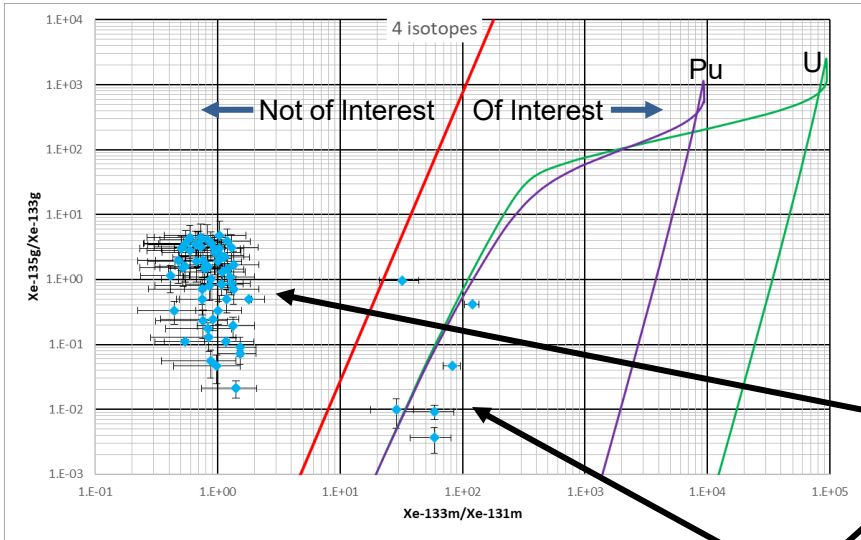
Example of Identifying Xe-133 Signal Injections in January 2014

	true positive rate	false positive rate
q50	0.924	0.115
q75	0.847	0.066
q90	0.784	0.036
q95	0.72	0.017
q96	0.716	0.015
q97	0.686	0.011
q98	0.657	0.008
q99	0.623	0.004

There is a tradeoff between true positives and false positives versus the quantile threshold.



# Identifying Xenon Anomalies in Measurements

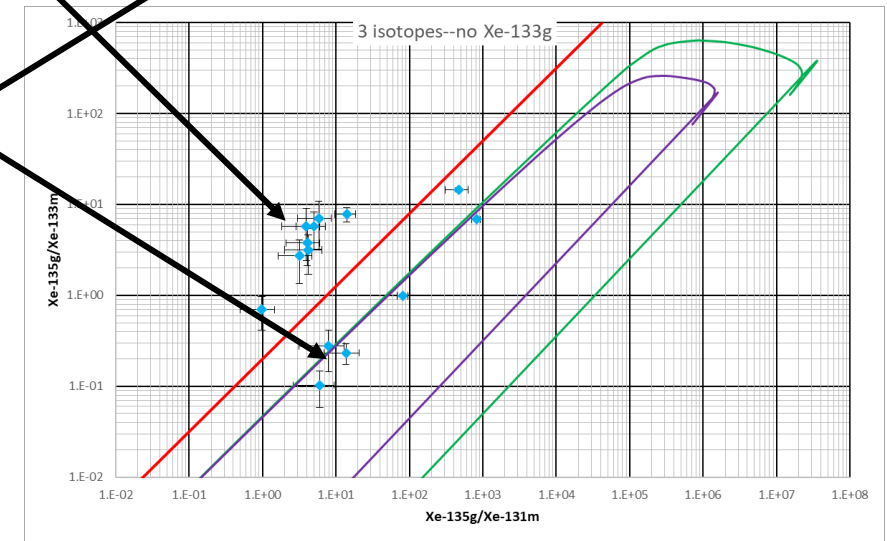
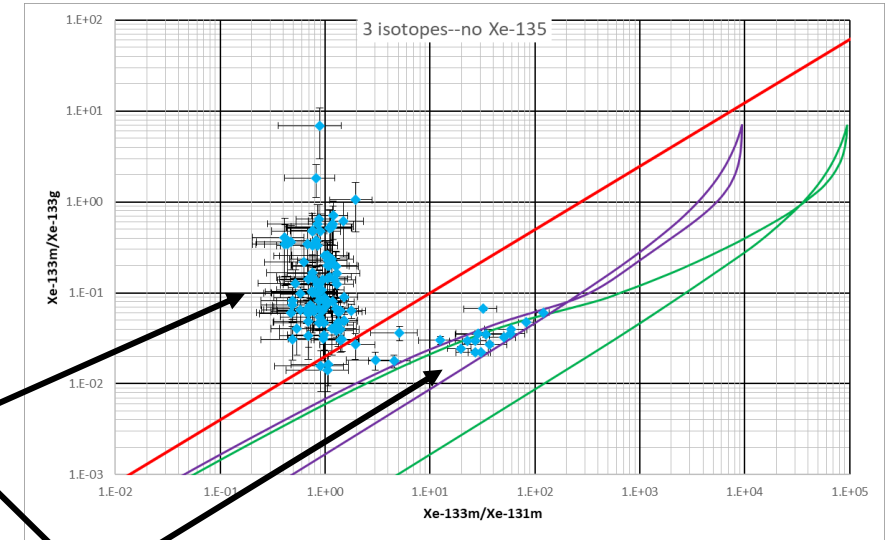


Xenon Ratios help identify anomalies

Xenon Samples not of interest

Xenon Samples of interest

Requires detecting multiple isotopes  
With fewer isotopes can use MDC



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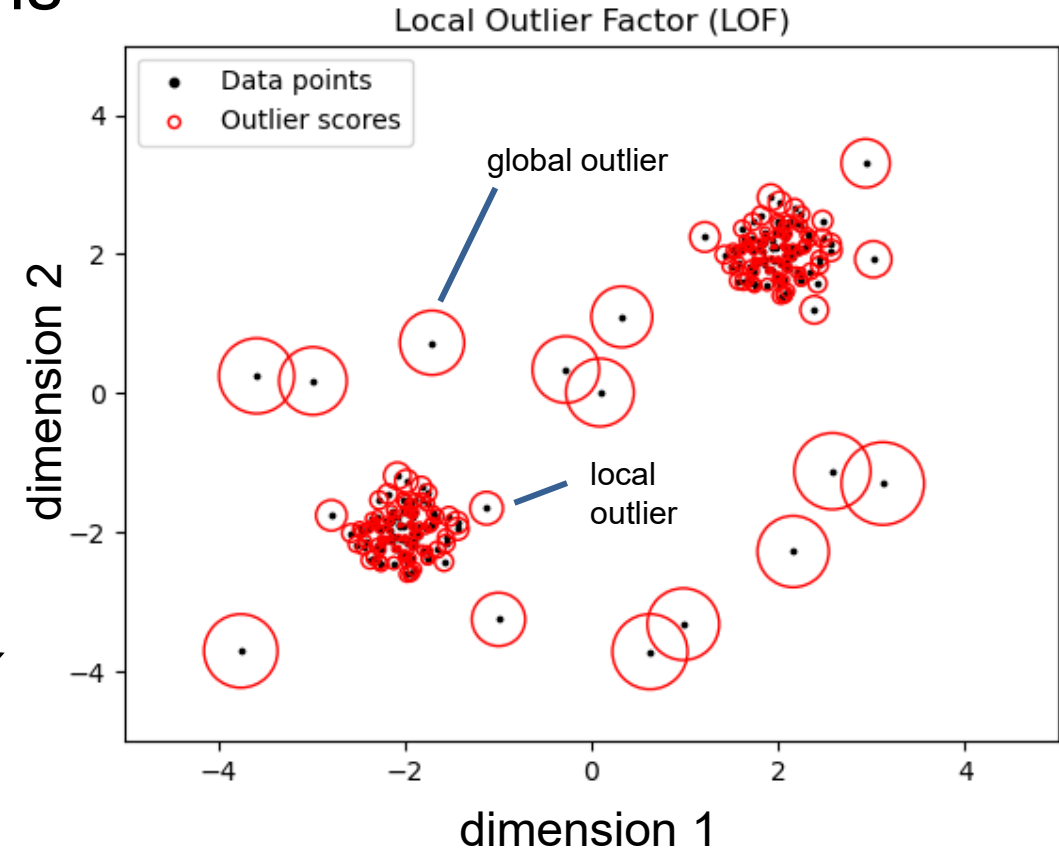
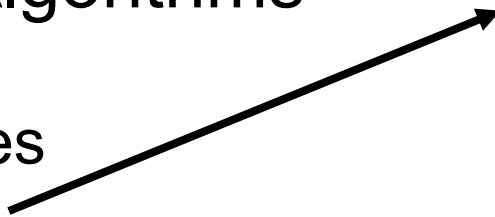


Image modified from 

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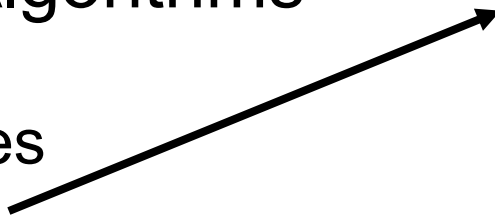
## Outlier/Novelty Detection Algorithms

Time series methods

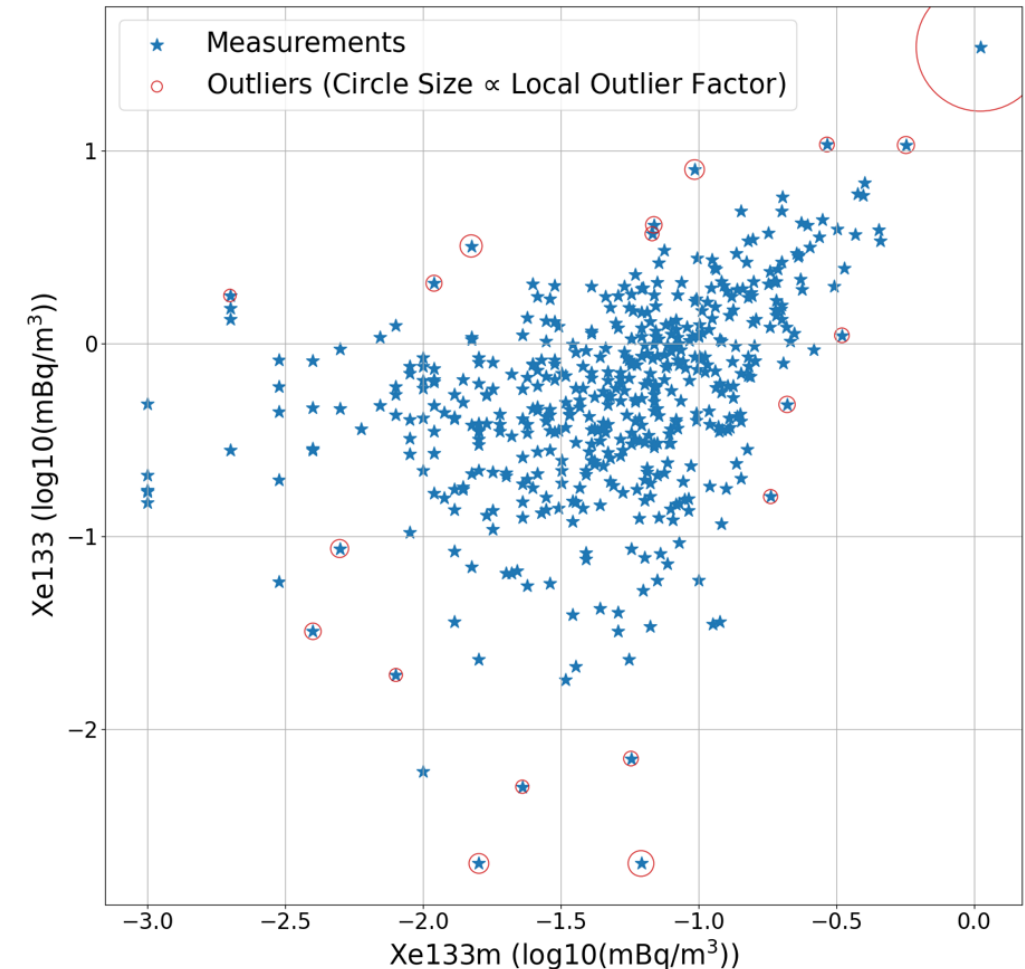
Machine learning approaches

Local Outlier Factor

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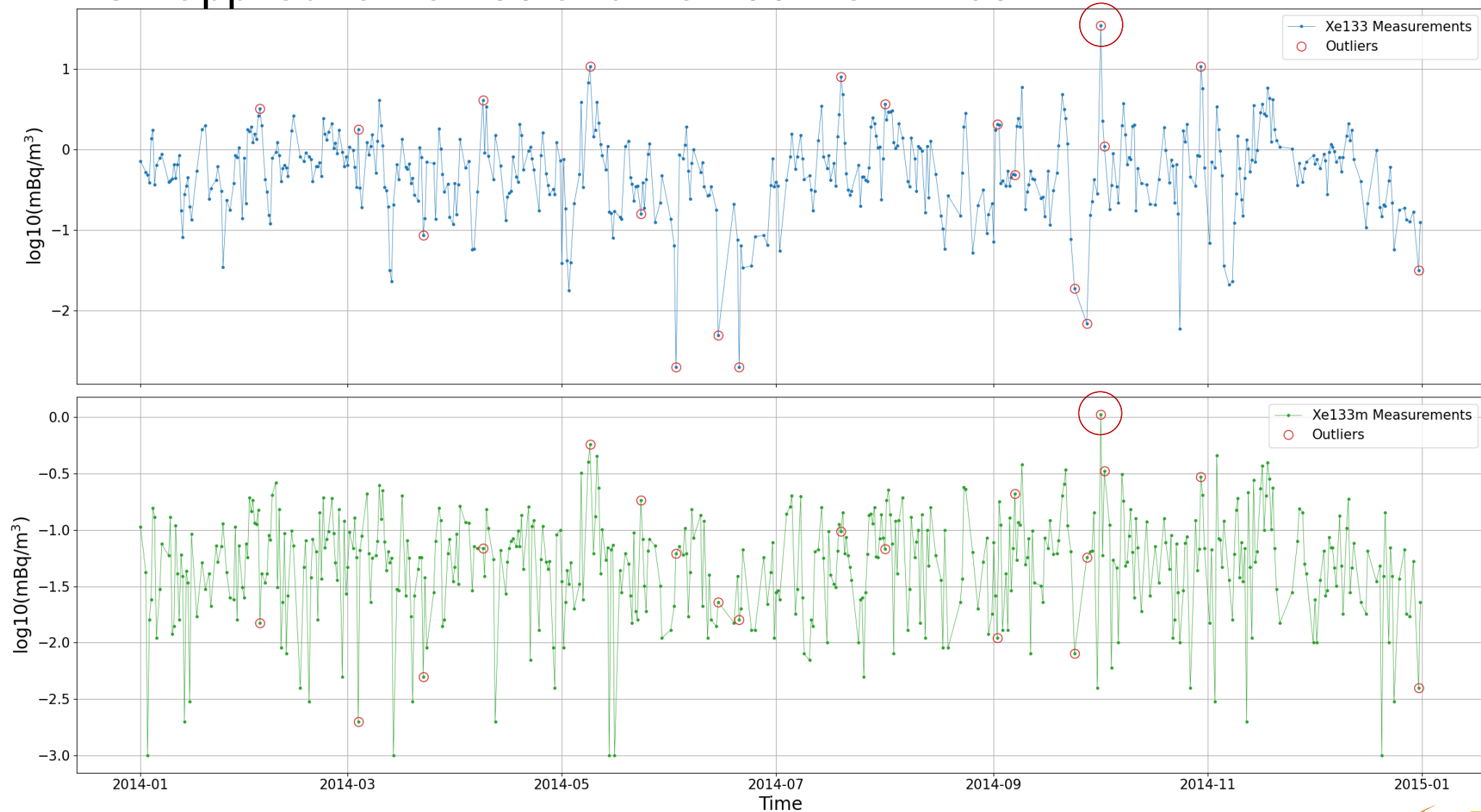


LOF applied to Xe-133 and Xe-133m at RN63 for 2014



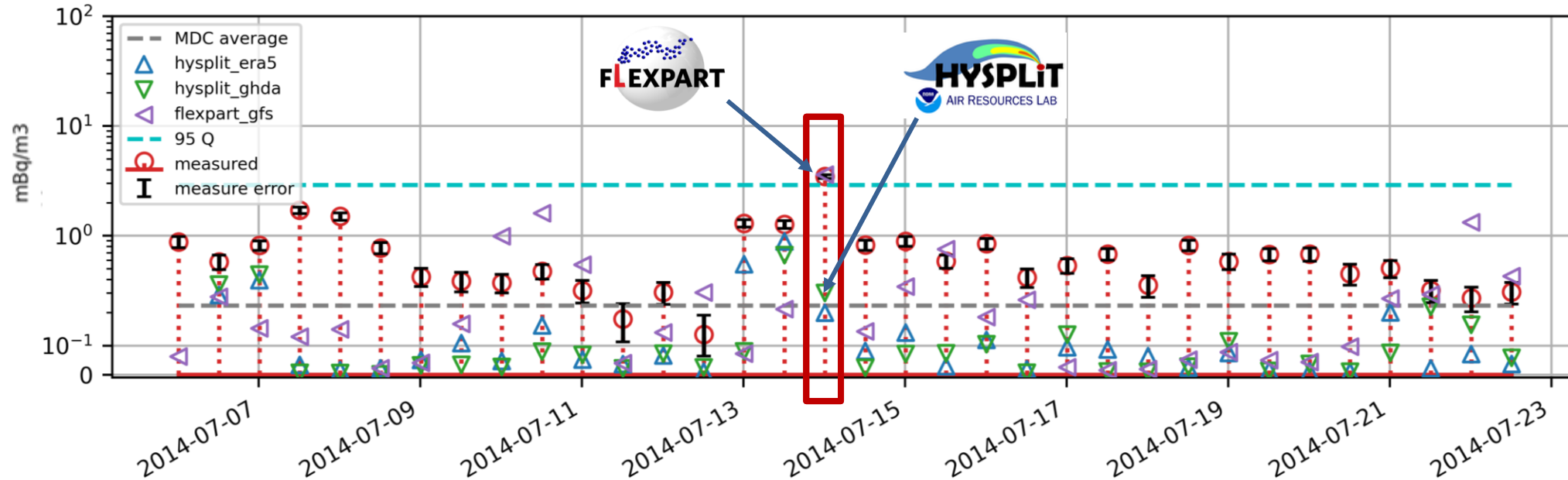
# Identifying Xenon Anomalies in Measurements

## LOF applied to Xe-133 and Xe-133m at RN63



# Identifying Anomalies With Measurements & Atmospheric Models

Xe-133 at Stockholm RN63



- Both models tend to underpredict Xe-133 during this period.
  - A case of low emissions or a bias in the atmospheric models?
- There was an elevated collection on 14 July at the 97<sup>th</sup> percentile.
- FLEXPART matches the elevation, HYSPLIT does not.
- Is the elevated collection an anomaly of interest?

# Identifying Anomalies With Measurements & Atmospheric Models

Regression methods can be used to combine ensembles of models, correct for model biases and errors, and provide predictions of IMS collections with uncertainty.



The diagram illustrates the relationship between the CTBTO International Monitoring System (IMS) and atmospheric models. On the left is the CTBTO logo, which includes the text "CTBTO PREPARATORY COMMISSION" and "INTERNATIONAL MONITORING SYSTEM". This is followed by an approximation symbol  $\approx$  and a function  $F$  in parentheses. Inside the parentheses are the logos for "HYSPLIT AIR RESOURCES LAB" and "FLEXPART". Below this visual representation is the mathematical equation: 
$$Y_{IMS} = F(Hysplit_1, Hysplit_2, \dots, Flexpart_1, Flexpart_2, \dots)$$

Train on data for previous periods → Apply to collections of interest

Other predictors can be incorporated, like collections from different IMS stations, environmental variables, categorization levels, etc.

# Identifying Anomalies With Measurements & Atmospheric Models

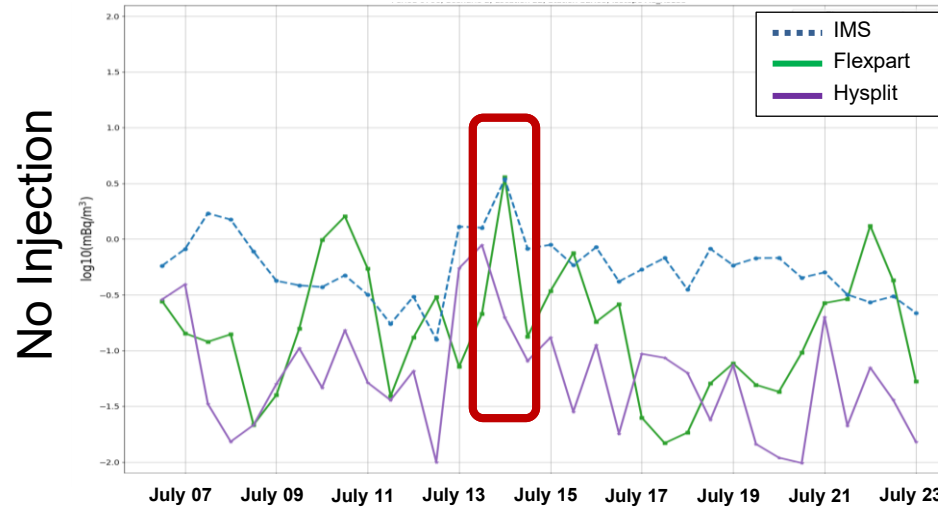
## Bayesian Ridge Regression for Xe-133 at RN63

Robust to outliers, easy to train, and provides uncertainty estimates.

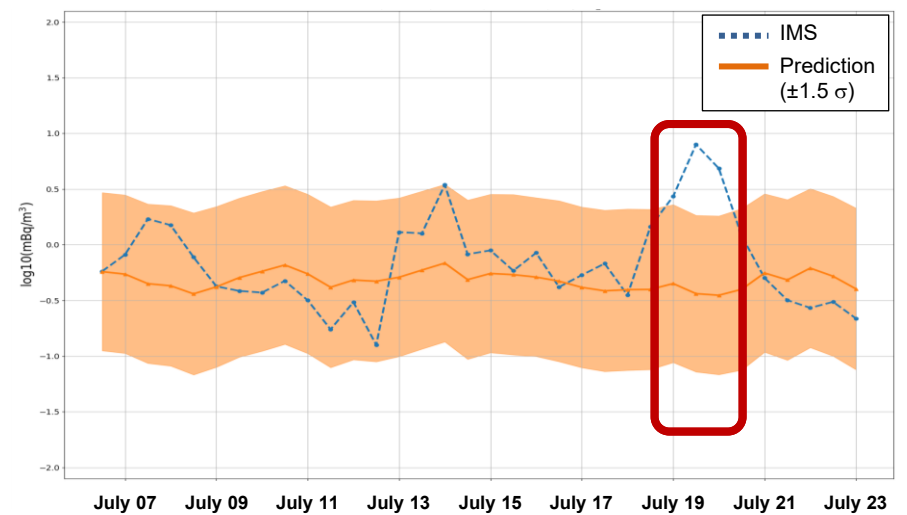
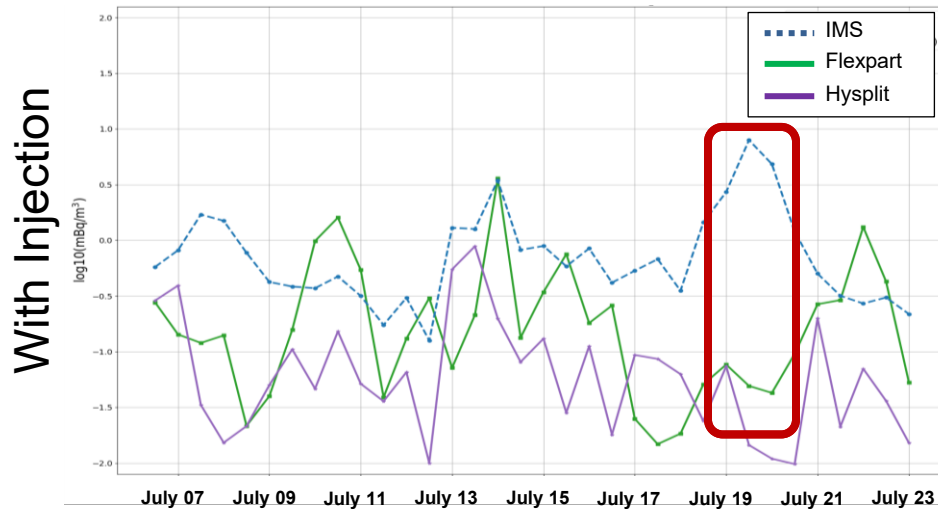
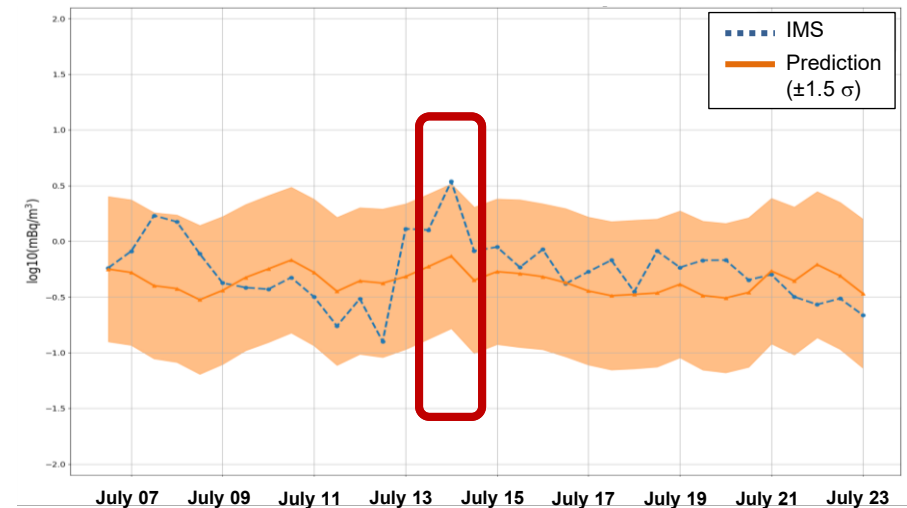
Elevated collection on 13-14 July lies within the regression prediction uncertainty.

Injected signal on 19-21 July is detected as an anomaly.

### Regression Inputs/Targets



### Regression Predictions



# Identifying the Origin of Anomalies

## Backwards Modeling

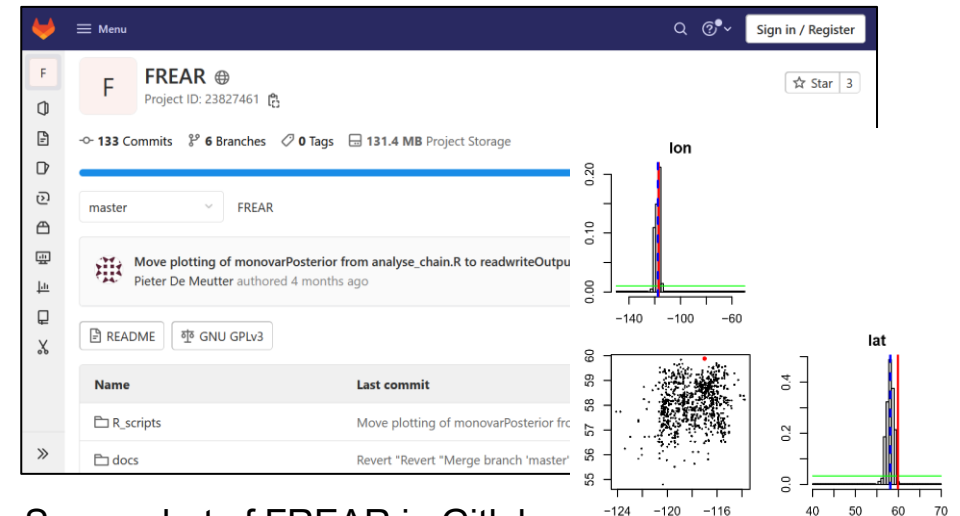
- Field of Regard (FOR)
- Possible Source Region (PSR)

## Probabilistic Methods

- *Forensic Radionuclide Event Analysis and Reconstruction Tool (FREAR)*
- *Machine Learning Approach*
  - Forward model runs are used to create synthetic detections/non-detections for training data and testing.
  - Once trained, millions of alternate source locations can be quickly evaluated.
  - Previously presented at WOSMIP and INGE.



3-day multi-model field of regard for JPX38 for collection for sample ID 2862643 using Web-Grape

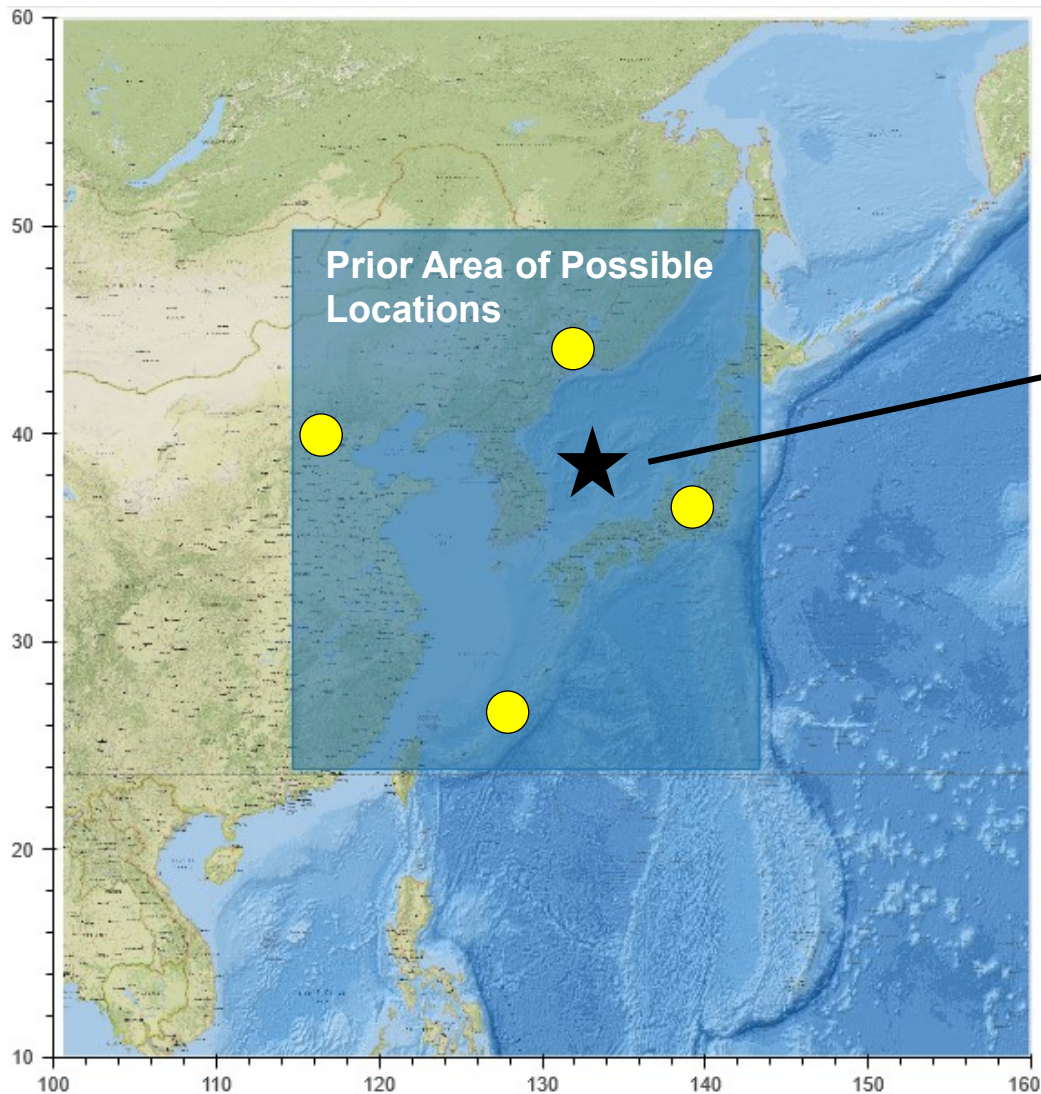


Screenshot of FREAR in Gitlab and test results.



# Identifying the Origin of Anomalies

## Machine Learning Example



Values to Generate Synthetic Observations

Latitude	Longitude	Start Hours	Duration Hours	Amount
38.1336	132.8962	65.4873	3.0234	896.9242

Forward  
Atmospheric  
Simulation



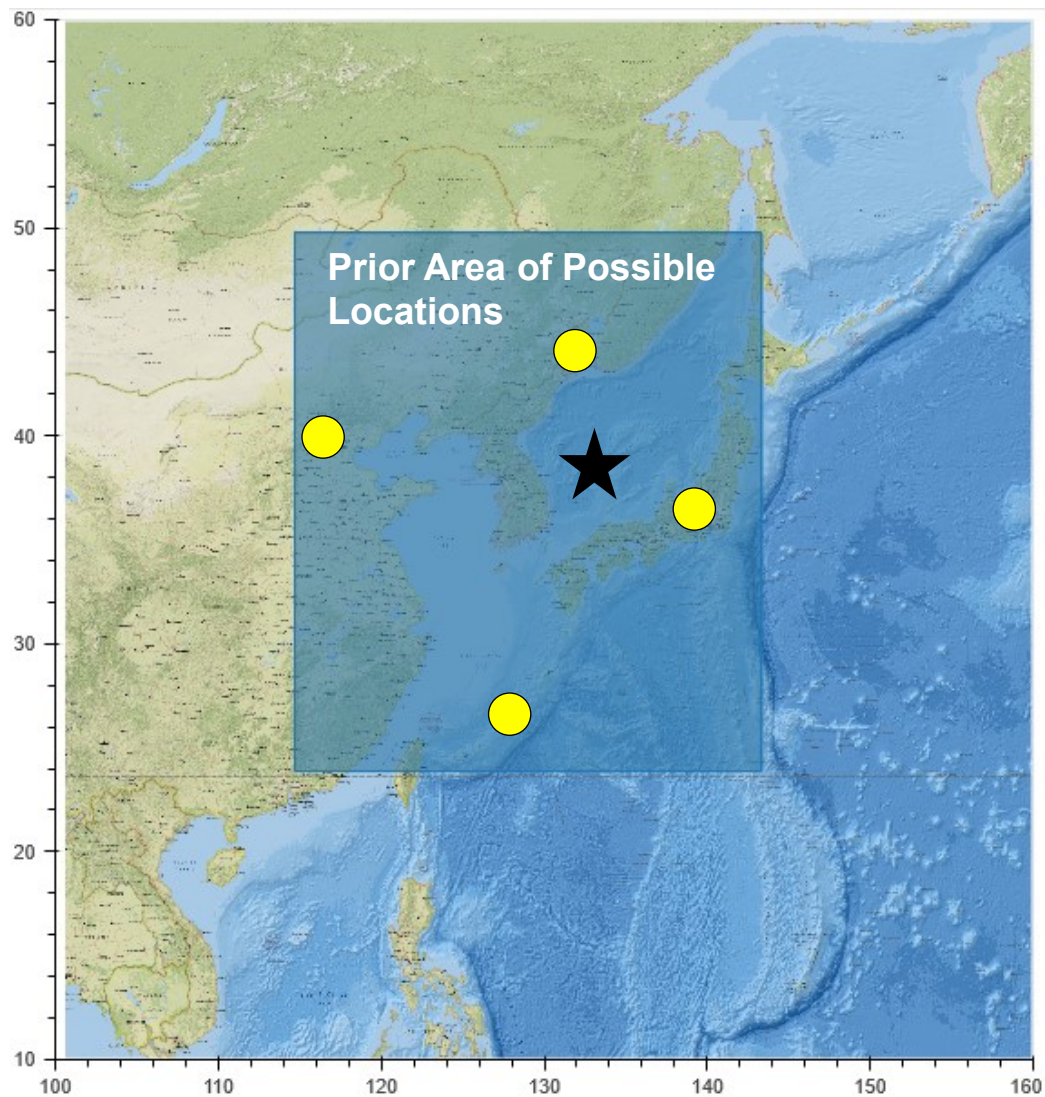
Machine  
Learning  
Backwards



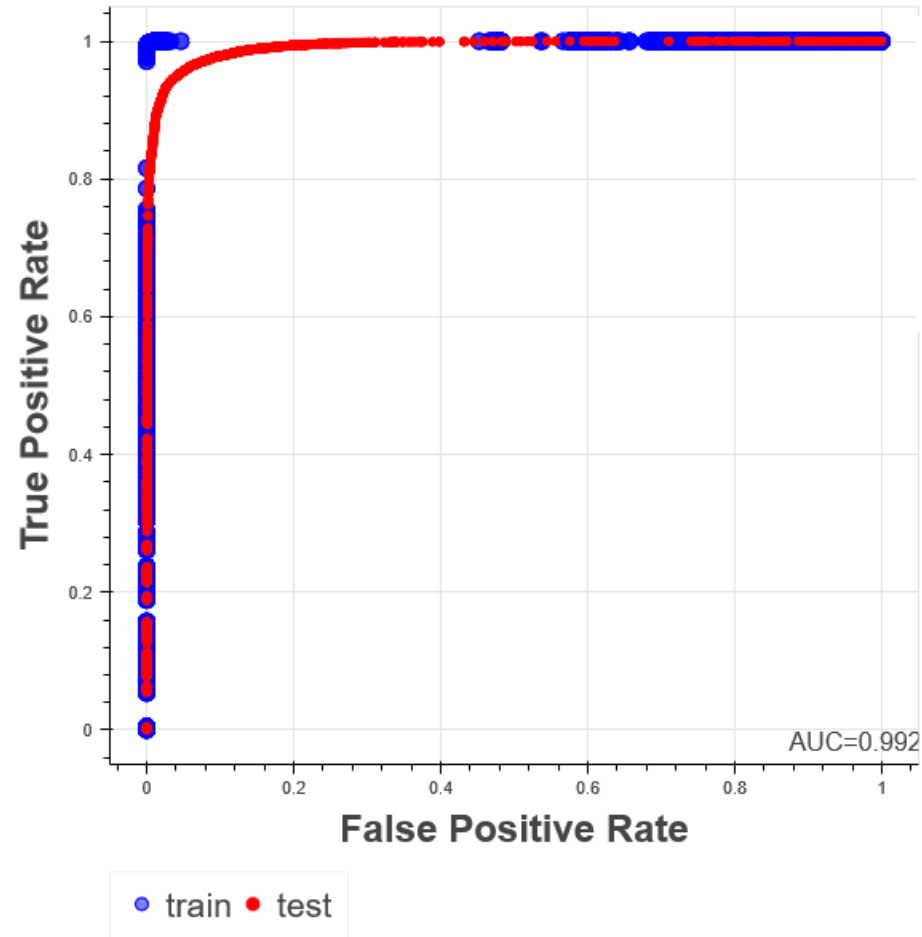
Site Name	Oct 09	Oct 10	Oct 11	Oct 12	Oct 13	Oct 14	Oct 15
RN20	0	0	0	0	0	0	0
RN37	0	0	0	0	0	0	1
RN38	0	0	0	0	1	1	1
RN58	0	0	0	0	0	0	0

■ = non-detect    ■ = detect

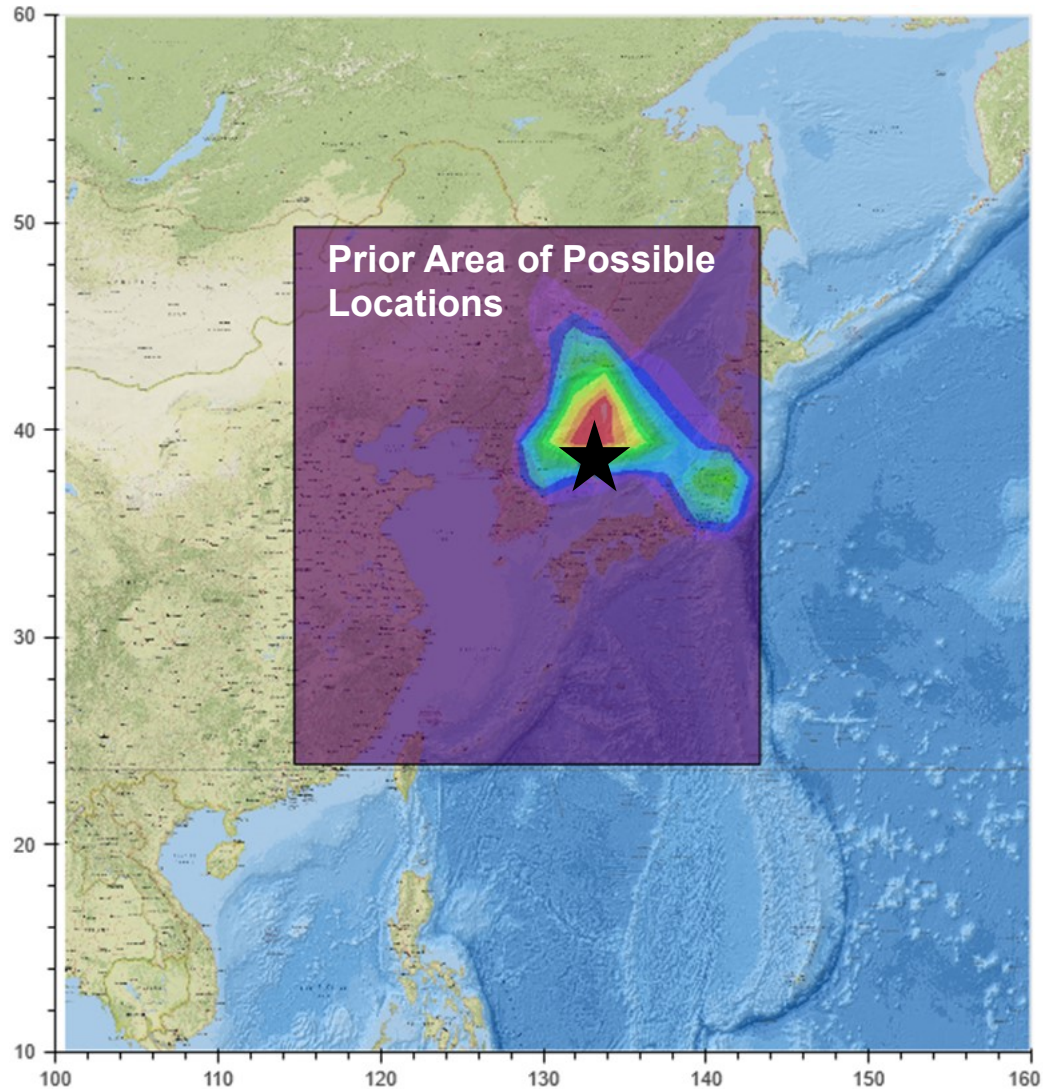
# Identifying the Origin of Anomalies



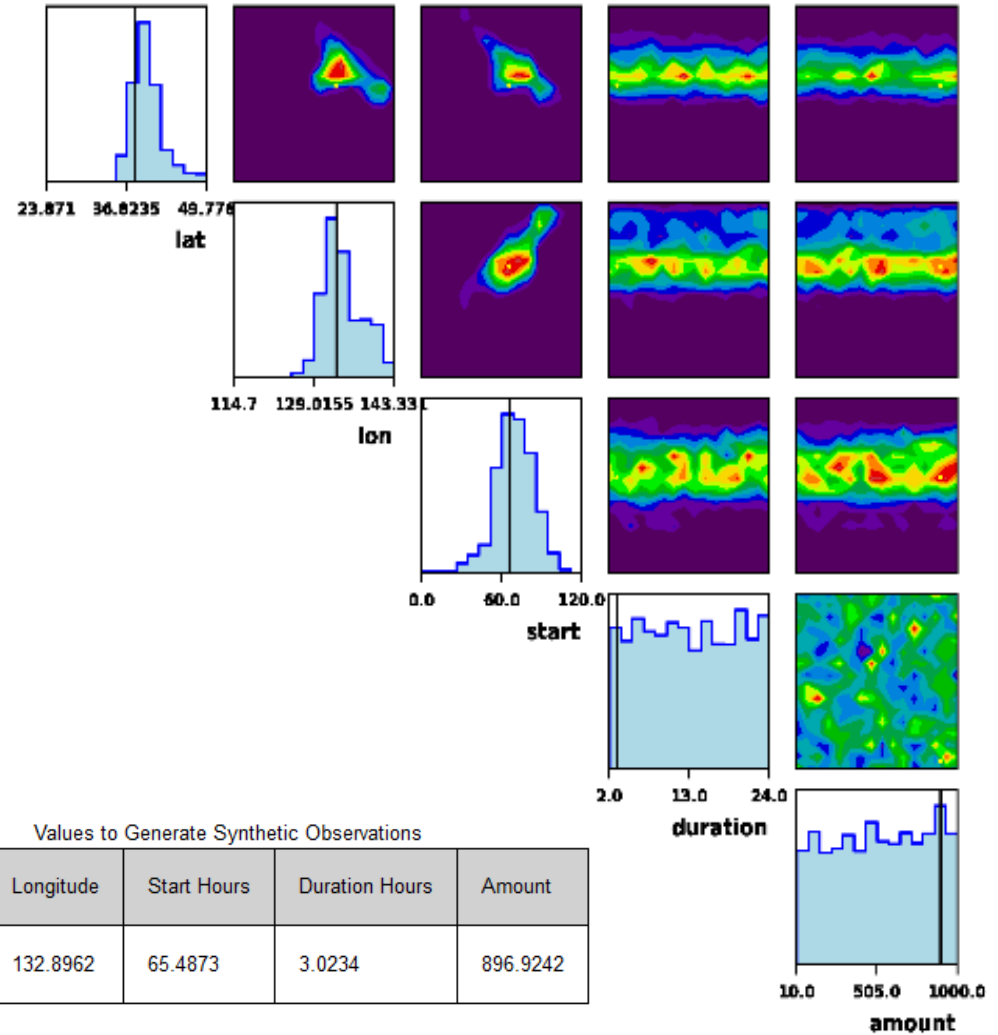
## Machine Learning Example



# Identifying the Origin of Anomalies



## Machine Learning Example



# Summary

- Xenon isotopes used for nuclear monitoring are highly variable in space and time due to
  - changes in weather.
  - the presence of many, widely distributed background industrial sources.
- Advanced methods using atmospheric modeling and statistical analysis are needed to
  - identify xenon anomalies of interest.
  - attribute the anomalies to background sources or nuclear testing.
  - determine the origin of detections.
- Through a collaborative effort, we
  - ran multiple atmospheric models to simulate xenon signals across the global IMS network in 2014.
  - developed and tested outlier and novelty detection methods using quantile approaches and unsupervised machine learning algorithms.
  - used supervised Bayesian regression algorithms to combine multi-model predictions and IMS collections for detecting anomalies with uncertainty.
  - applied probabilistic algorithms to locate the origin of suspected anomalies.

**Extra Slides**

# Atmospheric Models Can Be Used to Estimate Background Xenon

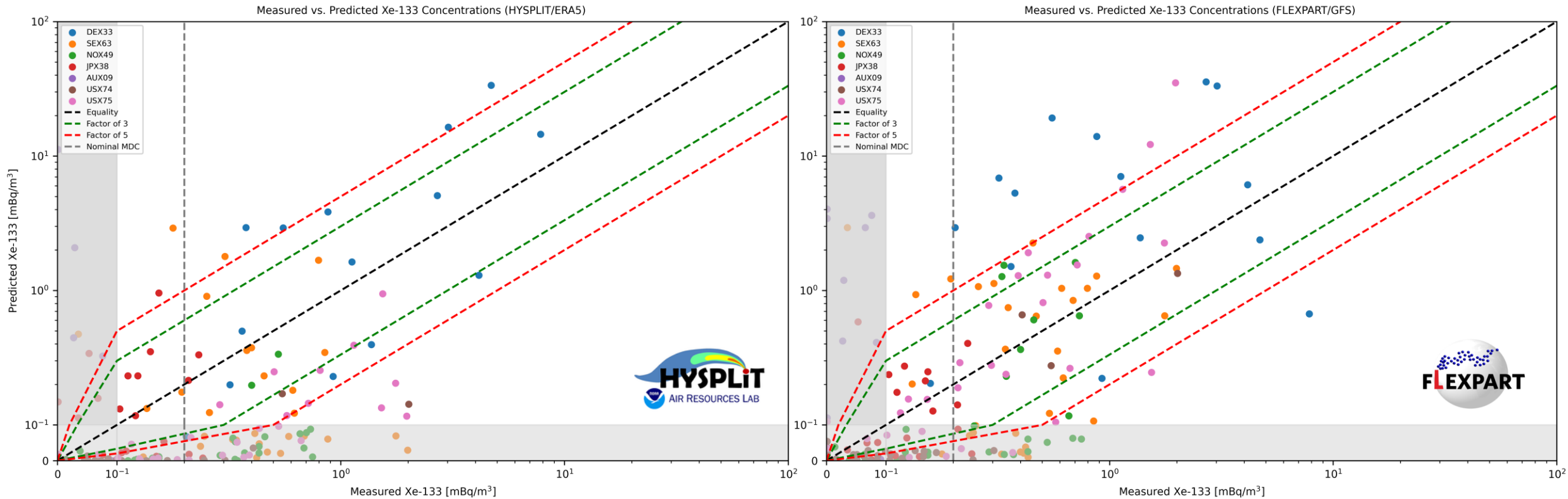
Atmospheric model performance statistics for January 2014

model	hysplit_era5			hysplit_gdas			flexpart_gfs		
	Factor 3x	Factor 5x	Corr	Factor 3x	Factor 5x	Corr	Factor 3x	Factor 5x	Corr
<b>NOX49</b>	0.1	0.133	0.45	0.033	0.1	0.208	0.267	0.367	0.389
<b>DEX33</b>	0.4	0.667	0.652	0.467	0.667	0.806	0.267	0.4	0.042
<b>SEX63</b>	0.267	0.367	-0.167	0.2	0.367	-0.097	0.433	0.667	0.109
<b>JPX38</b>	0.4	0.65	0.082	0.25	0.4	-0.448	0.65	0.65	0.12
<b>AUX09</b>	0.083	0.167	-0.257	0.167	0.25	-0.236	0	0.083	0.067
<b>USX74</b>	0	0.067	0.778	0	0.033	0.421	0.233	0.367	0.938
<b>USX75</b>	0.167	0.367	0.605	0.1	0.3	0.376	0.533	0.667	0.671
<b>23 stations</b>	<b>0.171</b>	<b>0.241</b>	<b>0.656</b>	<b>0.16</b>	<b>0.226</b>	<b>0.758</b>	<b>0.25</b>	<b>0.318</b>	<b>0.433</b>

Further data QC may be necessary (based on analysis of spectra)

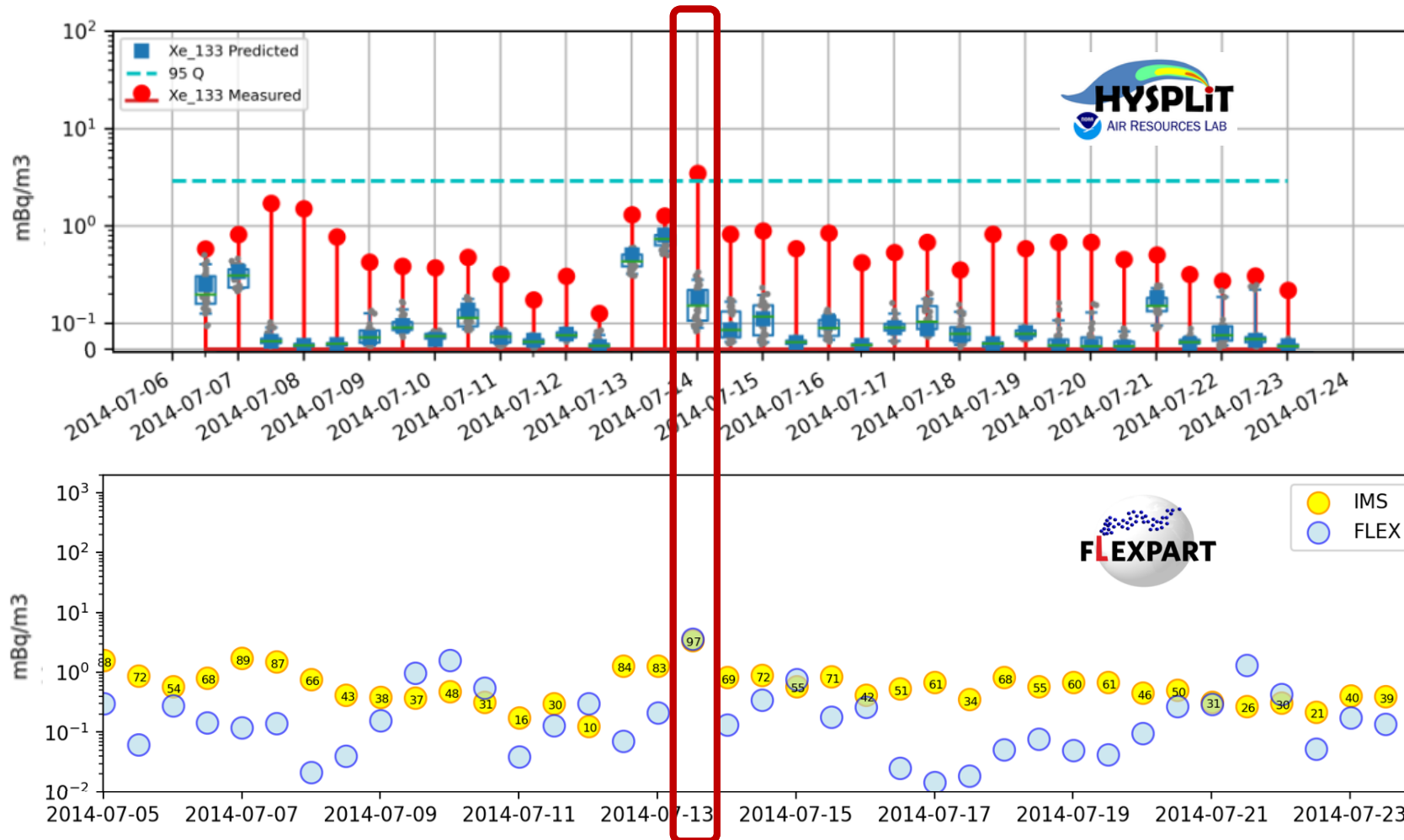
# Atmospheric Models Can Be Used to Estimate Background Xenon

## Atmospheric model performance for January 2014



# Identifying Anomalies With Measurements & Atmospheric Models

Xe-133 at Stockholm RN63



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- There was an elevated collection on 13-14 July at the 97<sup>th</sup> percentile.
- FLEXPART matches the elevation, HYSPLIT does not.
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# Identifying Anomalies With Measurements & Atmospheric Models

*Bayesian ridge regression performs well at many IMS stations*

Shown for July 2014 Xe-133 collections.

Performance evaluated by temporal correlation between predictions and independent collections.

Performance improves with increasing amounts of training data.

Some stations are more difficult to predict than others with atmospheric models.

